



# 2016 ExoMars/TGO

## Overview



- Joint ESA-NASA Mars Orbiter
  - Proposed launch in 2016
- Deliver ESA EDM
- 1 Mars year orbital science mission
- Telecom asset until 2022

### **Contributions**



### NASA / JPL Proposed Mission Elements



MAGIE



**EMCS** 

Focal Pare Electronics Trumon Ring Telescope Resident Branch Trumon Ring Telescope Resident Branch Trumon Ring Telescope Resident Resident



Electra

MATMOS



Atlas-V – Baseline JPL-JG1

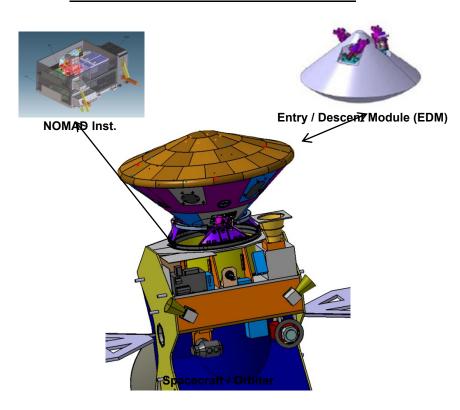


JPL DSN



JPL SRA

#### **ESA Mission Elements**





**ESA ESTRACK** 



**ESA ESOC** 

#### Slide 3

#### LV not selected yet Janis Graham, 5/18/2011 JPL-JG1

## 2016 ExoMars/TGO

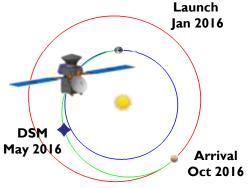
## EMTGO Mission Events (proposed)





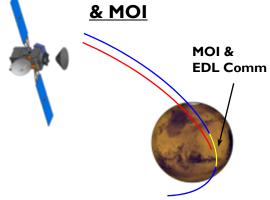
Atlas V 43 I

## CRUISE Launch



Type II Trajectory:  $C3 = 7.44 \text{ km}^2/\text{s}^2$ 

## & MOI



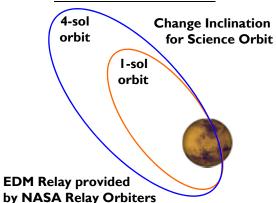
- EDM release at MOI 3 days
- Orbiter retargets to MOI altitude
- MOI captures to 4 sol orbit

## EDM RELAY & TRANSITION TO I-SOL ORBIT

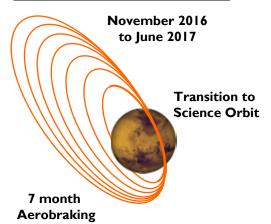
**EMTGO** in

launch

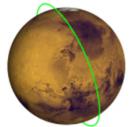
configuration



#### **AEROBRAKING PHASE**



## SCIENCE & DATA RELAY PHASE



- Science & Relay Orbit
- 400 km Frozen
- Rotates every 4 months
- Phased for 2018 Relay
- Science Phase: I Mars Year 6-2017 to 6-2019
- Relay Phase: 2018 Rovers Jan 2019
- Relay Phase: Future Missions through 2022

## Prioritized Science Objectives



- Detect a broad suite of atmospheric trace gases and key isotopologues.
- Characterize the spatial and temporal variation of methane and other species that could be signatures of active biological and/or geological processes (for example, C<sub>2</sub>H<sub>6</sub>, SO<sub>2</sub>, N<sub>2</sub>O) and of photochemical species that determine atmospheric lifetimes (e.g., representative O<sub>x</sub>, HO<sub>x</sub>, NO<sub>x</sub> species) and their source molecules (e.g., H<sub>2</sub>O).
- Localize the sources and derive the evolution of methane and other key species and their possible interactions, including interactions with atmospheric aerosols and how they are affected by the atmospheric state (temperature and distribution of major source gases; e.g. H<sub>2</sub>O).
- Image surface features possibly related to trace gas sources and sinks.

## Proposed Payload



#### **MATMOS**

Solar occultation Fourier transform IR spectrometer (w/ Canadian contribution)

#### **NOMAD**

Occultation + mapping IR, Vis, UV spectrometer (consortium of Belgium, Spain, Italy, UK)

#### **EMCS**

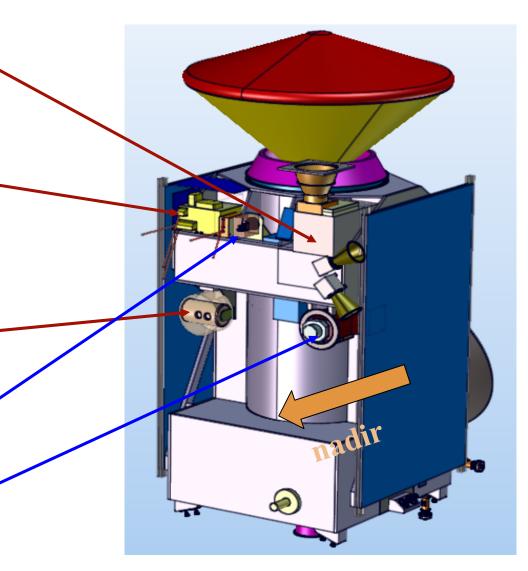
Thermal IR spectrometer

#### **MAGIE**

Wide-angle Vis-UV camera

#### **HiSCI**

High resolution, colour, stereo camera (w/ Swiss contribution)



#### Slide 6

#### **JPL-JG3** Is this all of the instruments?

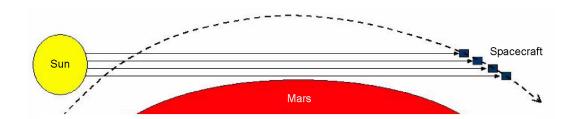
Janis Graham, 5/18/2011

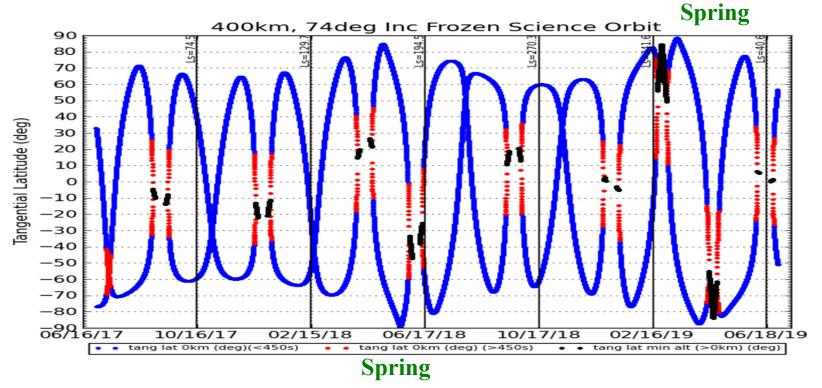
# 2016 ExoMars/TGO

### Solar occultation



- Ultrahigh sensitivity
  - Bright light source
  - Long pathlength
- Orbit inclination: 74°





## Nadir mapping



Ground track for 3 days

